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Impact of dust aerosols on the radiative budget, surface heat fluxes, heating rate profiles and convective activity over West Africa during March 2006

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Abstract

The present work analyzes the effect of dust aerosols on the surface and top of atmosphere radiative budget, surface temperature, sensible heat fluxes, atmospheric heating rate and convective activity over West Africa. The study is focused on the regional impact of a major dust event over the period of 9–13 March. Numerical simulations have been performed with the MesoNH model in which full interactions between radiation and dynamics are introduced, through various components representing size-resolved aerosol and cloud microphysics, radiative properties of particles and clouds, dynamics, and a surface model. Due to its importance on radiative budgets, a specific attention has been paid to the representation of dust SSA in MesoNH by using AERONET inversions. The radiative impacts are estimated using two parallel simulations, one including radiative effects of dust and the other without them. The simulations of dust aerosol impacts on the radiative budget indicate remarkable instantaneous decrease of shortwave (SW) radiations, with regional (09°–17° N, 10° W–20° E) mean of -160 W/m^2 during the 9 to 13 March period. The surface dimming resulting from the presence of dust is shown to cause important reduction of both surface temperature (up to 4°C over regions where high AODs occur) and sensible heat fluxes (up to 100 W/m^2), which is consistent with experimental observations performed over the same region. At the top of the atmosphere, the SW cooling (regional mean of -13.5 W/m^2) induced by mineral dust, although moderated by the longwave (LW) warming (regional mean of $+5 \text{ W/m}^2$), dominates the total net (shortwave+longwave) effect. The maximum SW heating occurs within the dusty layer with values comprised between 4 and 7°K by day and LW effect results in strong cooling (-6 to -16°K by day) below the dust layer. Finally, the simulations suggest the decrease of the convective available potential energy (CAPE) over the region in the presence of mineral dust.

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1 Introduction

Numerous studies have been dealing with estimation of direct effects associated with major aerosol types (urban/industrial, smoke and mineral dust) and related changes of the radiative budget at the surface, top of atmosphere (TOA) and within the aerosol layer. Such studies (Ramanathan et al., 2001; Lelieveld et al., 2002; Huebert et al., 2003; Tanré et al., 2003; Abel et al., 2005; Roger et al., 2006; Ramanathan et al., 2007, Haywood et al., 2008) underline that fine aerosols (urban/industrial and/or smoke particles) decrease significantly surface incoming shortwave radiations and generally increase (or rarely decrease in some specific cases, Haywood and Shine, 1995) outgoing shortwave fluxes reflected back to space. Except in cases of pure scattering particles, the net effect for the atmosphere is positive revealing a gain of solar energy within the aerosol layer. The effect of dust particles is more complex because of their ability to interact both in shortwave and longwave radiations. As a result, top of the atmosphere SW and LW forcings are generally opposite: the presence of the dust generally increases (except above high surface albedo) upward fluxes in shortwave and decreases them in longwave spectral range.

Over a given region, these various radiative effects, occurring at different atmospheric levels (surface, TOA and within the aerosol layer), all contribute to the climatic response. Thus it is very important to understand how these forcings can act together and alter the surface energy budget (heat fluxes), cloud properties (semi-direct effect), atmospheric dynamics as well as hydrological cycle (modification of precipitation regimes). For example, the studies conducted during the INDOEX experiment (Ramanathan et al., 2001) over the Indo-Asian region reveal that surface cooling and tropospheric heating associated with particles are able to significantly perturb the regional climate and tropical rainfall patterns with implications for global climate (Ramanathan et al., 2001).

Perturbations similar to those observed during INDOEX should also be expected over West Africa, which is subject to large smoke and mineral dust loadings. The

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Sahara desert acts as the strongest source of mineral dust aerosol in the world (Woodward, 2001). The most intense dust events in Sahara and Sahelian regions are frequently generated in spring (Marticorena et al., 1997) due to favourable large scale conditions, but there is also a significant dust production throughout the year. In addition, the burning of agricultural waste in the Sahelian region during the dry season makes of West Africa one of the strongest source of biomass burning aerosols.

Presently, the potential impact of dust and smoke particles on the West Africa climate is unclear and the estimation of this impact is one of the main purposes of the AMMA (African Monsoon Multidisciplinary Analysis) program (Redelsperger, 2006), a major international campaign investigating different aspects of the African monsoon (climate dynamics, hydrological cycle, aerosol/chemistry and impacts). Ground-based measurements, coupled with aircraft and satellite observations of aerosols have been developed within the AMMA program for investigating their influence on the radiation balance of the earth/atmosphere system.

The present study investigates the direct radiative forcing of mineral dust and its potential impact in terms of change in surface temperature and sensible heat fluxes, heating rate profiles and atmospheric dynamic. This work has been conducted by using the MesoNH meso-scale forecasting model (Lafore et al., 1998) fully coupled with a dust production and transport model (Grini et al., 2006), a radiation scheme and an explicit land surface model. This kind of “on-line” model enables to investigate how the aerosols direct radiative forcing impact the surface-atmosphere system. Due to the significant impact of dust SSA on the radiative budget (both at TOA, surface SRF, and into the dust layer) and heating rate profiles, we used dust refractive indexes retrieved by AERONET (Dubovik and King, 2000) together with the size distribution for calculating the shortwave dust SSA used in our MesoNH simulations. In parallel, experimental observations obtained at Djougou (Northern Benin) allowed to complement modelling simulations. Here, we perform a case study for an intense dust event occurring during March 2006 (Slingo et al., 2006; Milton et al., 2008; Tulet et al., 2008), leading to dust aerosol optical depth (AOD) larger than 2 (at 550 nm) over a large part of West Africa.

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The structure of the paper is the following. The first part (Sect. 2) briefly describes the dust outbreak investigated in this work. Section 3 describes the experimental observations performed at Djougou (Northern Benin), especially in terms of surface sensible fluxes. Section 4 presents the developments of the dust optical properties in MesoNH.

5 Finally, the impact of dust aerosols on the radiative budget, surface heat fluxes, heating rate profiles as well as convective activity is discussed in the Sect. 5. The conclusions and summary of this work are given in the Sect. 6.

2 Model simulations of the March 2006 dust outbreak

The simulations begin at 00:00 UTC on 7 March 2006, and end at 00:00 UTC on 14
10 March 2006. Two-way nested grid domains have been used. The large domain (36-km resolution) between 3.1° S and 31.7° N in latitude and 25.64° W and 35.64° E in longitude, gives a large scale synoptic view of west Africa. The embedded smallest domain (12-km resolution) centered on the north-west of Nigeria covers a large part of the AMMA campaign area (latitudes 4.3° N and 17.6° N and longitudes 4.19° W and
15 16.24° E). The vertical resolution is composed of 60 stretched vertical levels reaching the altitude of 34 000 m, whereas 30 levels are located in the boundary layer between the surface and 2000 m. Initialization and lateral boundary conditions of the large domain were taken from the ECMWF analysis.

The dust AOD evolution (at 550 nm) as simulated by MesoNH is fully described in Tulet et al. (2008). We just report here (Fig. 1) the simulated dust AOD for the 9,
20 10, 11 and 12 March (at noon). Briefly, results show that a strong belt of high AOD appears from Chad to Senegal after two days of simulation. Different AOD maxima have been simulated in the central part of Lybia (3.2), Chad (3.4), the Darfour area (2.6), the south of Niger and the north of Nigeria (3.2). This high AOD belt is consistent
25 with MSG-SEVIRI and MODIS-AQUA images (Tulet et al., 2008). On March 10, the plume of dust AOD spread to the south, reaching the Gulf of Guinea. In particular, three intense AOD maxima exceeding 3.2 have been simulated around Nigeria (from

Benin to southern Chad and Cameroon). On March, 12, the intense dust plume was advected southward to the Guinea Sea and the Atlantic Ocean. The AOD decreased on the whole domain, but there were still significant peaks exceeding 1.6 from Nigeria to Ghana, 2 over Guinea, and 2.6 over the Guinea Gulf.

3 Data analysis of sensible heat fluxes

Associated with ground-based observations of microphysical, chemical and optical properties of aerosols (Pelon et al., 2008; Mallet et al., 2008), experimental measurements of surface sensible heat fluxes (H) were conducted at Djougou. Here, we just detail the methodology used for estimating surface fluxes, based on the eddy covariance method, which is considered as the reference method for vertical flux scalar measurement (Beverland et al., 1996).

This method is based on the direct high frequency measurement of the two components of the vertical flux of a scalar: the vertical wind speed “w” and the scalar itself, “c” here, which can be temperature, water vapor mixing ratio or CO₂ concentration. Flux comes as the integral of the product of the vertical wind speed fluctuation w’ and of the scalar c’:

$$F_c = \overline{w'c'} = \frac{1}{T} \int_0^T w'(t)c'(t)dt = \frac{1}{T} f(t) \quad \text{with} \quad f(t) = \int_0^T w'(t)c'(t)dt \quad (1)$$

Although this method presents a number of advantages, the processing of raw data shows that data control is necessary to guarantee their quality for further use (Foken et al., 1996; Affre et al., 2000). Control of the turbulence stationarity, a fundamental hypothesis to be fulfilled when determining turbulent fluxes, was performed here. It is relatively simple step in processing data since all the turbulence functions involved are available and can be verified at all post-treatment steps (Mann and Lenschow, 1994).

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Sensible, latent heat and CO₂ fluxes were measured together with the Eddy Correlation (EC) technique for a period extending from 1 September 2005 to 6 April 2007. Experimental set up included a Gill[®] R1 3-D sonic anemometer, a Licor[®] 7500 open path Infrared Gas analyzer, and a lab-made datalogger. 8 Hz logging of horizontal wind components (U , V), vertical wind component (W), temperature (T), H₂O and CO₂ data was performed on a 30 min basis. Sonic and Licor 7500 path were located at 8 m above ground level at the tip of a 2 m pole born by a scaffolding.

4 Radiative effect parametrization and dust optical properties

The MesoNH model uses the radiation code of ECMWF (Fouquart and Bonnel, 1980; Morcrette and Fouquart, 1986), which computes the radiative fluxes of solar and thermal infrared radiation. We used here the standard formulation of absorptivity/emissivity of longwave radiation for aerosols in the ECMWF model. Scattering of longwave radiation by mineral dust is neglected in our radiative calculations. As reported by Dufresne et al. (2002), this may lead to an underestimation of longwave radiative forcing by up to 50% at TOA and 15% at the surface.

Here, we discuss about the simulations of dust SSA. The simulations of dust AOD were described in Tulet et al. (2008). The aerosol SSA is the optical property that determines the sign of TOA forcing (Lio and Seinfeld, 1998), depending on a critical surface albedo ($R_{s,c}$) (Fraser and Kaufman, 1985). As an example, for a SSA of 0.90, $R_{s,c}$ is around 0.3. As the West African continent is characterized by a significant North-South gradient of the surface albedo, a rigorous estimation of the dust SSA is necessary for well exploring the mineral dust impact at the regional scale.

The calculations of SSA requires the knowledge of the size and the complex refractive index ($m=n-ik$, where n and k are the real and imaginary part, respectively) of the particle. While the size distribution is usually provided by aerosol model output, the refractive indexes of the dust are usually fixed in climate simulations to the values from available databases (e.g. Koepke et al., 1997, Hess et al., 1998). However, the recent

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findings indicated a possibility of strong deviations of these literature values from the real ones, which are controlled by dust chemical composition (especially percentage of hematite) and mixing with other primary (black carbon) and/or secondary aerosols (organics, sulphate, . . .). For example, the remote sensing study of Dubovik et al. (2002) reported much lower desert dust absorption in visible spectral range than commonly used in aerosol properties databases. The reported difference in dust SSA could lead to substantial errors on dust radiative forcings.

Therefore, in order to take into account the most realistic dust SSA in our numerical simulations, we have used dust size distributions as simulated by MesoNH (Tulet et al., 2005; Grini et al., 2006) together with dust refractive indexes retrieved from AERONET/PHOTONS radiometers measurements (Dubovik and King., 2000) during the dust outbreak. Our methodology is described in the Fig. 2. Dust size distribution and refractive index have been used in Mie calculations to derive dust optical properties required for radiative forcing calculations (extinction coefficients, SSA and g). As reported by Milton et al. (2008), the assumption of spherical particles appears here reasonable. Indeed, author's performed optical calculations by using mixtures of oblate and prolate spheroids and concluded that difference in SSA and g are lower than 10%.

During the dust outbreak, the mean values (and associated errors) of the refractive index (at 440 nm) are about 1.45 ± 0.04 and 0.0029 ± 0.0014 , for the real and imaginary parts, respectively. Figure 3 displays bulk dust SSA at 560 nm as obtained (at noon) over the West Africa region at the surface and 2 km height for the 9 and 10 March, showing SSA about 0.90–0.92 in the dust plume. Lower values are observed near sources, where the contribution of larger particles, which are more absorbing than fine dust in the SW, is maximum.

Such SSA values are found to be lower than those recently obtained over North Africa during DODO1 and DODO2 (0.99 and 0.98) (McConnell et al., 2008), DABEX during January 2006 at Niamey (0.99) (Osborne et al., 2008) and SHADE (0.97) (Haywood et al., 2003). This is due to the fact that the ORILAM scheme takes into account the coarse mode of the dust size distribution while both campaigns mentioned account-

ing only for the accumulation mode. McConnell et al. (2008) reported that the addition of the coarse mode induces a substantial change on SSA (at 550 nm) during DODO2 from 0.98 to 0.90, underlying that an accurate measurement of the coarse mode in mineral dust is extremely important for investigating dust radiative effects and climatic impacts.

The dust SSA spectral dependence was also rigorously included in our calculations, indicating that SSA increases with wavelengths in the dusty layer, changing from 0.88 to 0.91 (10 March), at 360 and 560 nm, respectively. This result is quite consistent with two years (1998–2000) of AERONET observations for mineral dust over Persian Gulf and Saudi Arabia, revealing changes of dust SSA (estimated over the total atmospheric column) from 0.92 to 0.96, at 440 and 670 nm, respectively (Dubovik et al., 2002).

5 Results and discussion

5.1 Dust effect on surface downward radiation

In this section, we discuss the effect of dust on surface (SRF) downward radiation, both in shortwave (SW) and longwave (LW) spectral ranges. We characterise this effect by calculating the flux difference at the surface:

$$\text{SRF_SW} = (\text{SW_DOWN})_{\text{DUST}} - (\text{SW_DOWN})_{\text{NONE}}. \quad (2)$$

where SW_DOWN is the downward shortwave radiative fluxes (in W/m^2). The same convention is used in the longwave region for calculating SRF_LW. Since atmospheric and surface variables are perturbed by dust aerosol in the DUST simulation, SRF_SW (and SRF_LW) cannot be not strictly defined as a “radiative forcing” (Haywood and Boucher, 2000) and it rather represents the overall effect of dust and atmospheric feedback on surface incoming radiation. With this convention, a negative sign of SRF_SW (or SRF_LW) implies an overall cooling effect at the surface and a positive one a heating effect. As reported by Roger et al. (2006) and Mallet et al. (2008), we consider that

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the radiative effect of aerosols at the surface is defined with an accuracy of 10%, due to uncertainties on SSA, asymmetry parameter and surface albedo.

5.1.1 Solar spectral range

Figure 4 reports the instantaneous shortwave surface effect (simulated at noon) for the 9, 10, 11 and 12 March. Results presented in the Fig. 4 clearly indicate that solar radiative fluxes are significantly reduced at the surface when the radiative effect of dust is included in MesoNH simulations. The reduction of the surface illumination is due to the reflection of solar radiations on the dusty layer but also because of absorption by mineral aerosols as dust SSA is lower than 1 (see Sect. 4).

As shown in the Fig. 4, SRF_SW (at noon) is almost everywhere negative with values comprised between -200 W/m^2 and quite remarkable maxima of -400 W/m^2 . In addition to continental regions, significant effects are also observed over oceanic zones as Atlantic ocean ($\approx -200/-300 \text{ W/m}^2$) and Gulf of Guinea ($\approx -200/-300 \text{ W/m}^2$), especially for the 11 and 12 March, when dusts are advected southward of the domain. It should be outlined here that surface differences are “instantaneous” (not averaged for a day) and reported for a specific time (here at noon). Our results of simulations are shown to be quite consistent with downward solar fluxes as measured at Djougou (Fig. 5), showing a large decrease in solar downward fluxes (around -150 W/m^2) during the dust outbreak (9–13 March).

Such a decrease has been also observed by Slingo et al. (2006) who reported that the dust storm dropped by around 250 W/m^2 the incoming solar fluxes at the surface at Niamey. Milton et al. (2008) modelled that dust particles considerably reduce net downward shortwave flux at the surface over a large part of West Africa with maximum of 200 W/m^2 . In addition, in the framework of the AMMA program, Derimian et al. (2008) estimated the aerosol surface radiative forcing over M’Bour, Senegal. The estimations were based on the aerosol properties derived from AERONET retrievals and validated against broad-band radiation measurements conducted during AMMA campaign period at M’Bour observational site. The instantaneous surface radiative forcing

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obtained for 10 March and for the midday time was of about 220 W/m^2 , which is close to our simulation results. Direct comparisons (Fig. 6) between downward fluxes as simulated by MesoNH and observed from pyranometers at Djougou for the 10–12 March period (pure clear-sky days) show a relatively good agreement.

Over West Africa, our results are found to be comparable with those ($-500 < \text{SRF}_{\text{SW}} < -200 \text{ W/m}^2$) reported by Grini et al. (2006), who studied a dust outbreak occurring over Mauritania and Senegal by using the MesoNH model. Helmert et al. (2007) reported also values comprised between -200 and -500 W/m^2 (simulated at noon) during a Saharan dust outbreak occurring in October 2001 by using the chemistry-transport model MUSCAT with a dust emission scheme.

Figure 4 displays some isolated positive values, with magnitudes comprised between $+100$ and $+200 \text{ W/m}^2$, over Ghana (9 March) and Southern Nigeria (above the Gulf of Guinea, for the 10 March), implying an increase of solar radiations reaching the surface when dust are included in simulations. As our calculations account for the atmospheric feedback, these changes are due to local modifications of cloud cover. On 10 March, a significant modification (not shown) in cloud cover appears between the control simulation (without dust radiative effect), where low clouds are developed over the Gulf of Guinea, and the second one (including dust effect), where clouds are dissipating. At this time, we do not know the processes responsible for modifying cloud cover over the region, and more detailed studies on this specific point has to be performed.

Averaged at the regional scale ($09\text{--}17^\circ \text{ N}/10^\circ \text{ W}\text{--}20^\circ \text{ E}$), our simulations indicate SRF_{SW} (obtained for the 9, 10, 11, 12 and 13 March, see Table 1) comprised between -108 and -207 W/m^2 (for dust AOD ranging from 0.60 and 1.23). These results clearly underline that dust aerosols produce a strong reduction of the incoming shortwave surface flux over a large part of the Western African region.

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5.1.2 Infrared spectral range

We present here the dust surface effect occurring now in the infrared spectral range (SRF_LW). The convention used is similar to the one (relation 2) used for solar wavelengths. Before presenting our results, we remind here that MesoNH simulations could underestimate the surface longwave effect by up to 15% (Dufresne et al., 2002), due to the fact that the scattering of longwave radiation is neglected.

In regional average, SRF_LW simulated for the 9, 10, 11 and 12 March (at noon) indicate values comprised between -3 and -5 W/m^2 (Table 1). The sign of the surface LW difference is negative and opposite to observed and/or simulated radiative effect at Niamey, where Slingo et al. (2006) reported that the downward thermal emission from the atmosphere display a significant peak on 7 March during the same dust outbreak. By using transfer radiative models, Liao and Seinfeld (1998) reported diurnal longwave surface forcing comprised between $+0.9$ and $+1.4 \text{ W/m}^2$, depending of the dust layer altitude. Dufresne et al. (2002) indicate also positive surface forcings comprised between $+21$ and $+29 \text{ W/m}^2$, for six different atmospheric profiles. Finally, three dimensional modelling have been performed by Yoshioka et al. (2007), who reported a mean annual value of $+14 \text{ W/m}^2$ over the North Africa region, by using the CCSM3 model. In our simulation, the average decrease of downward LW radiation in surface can be explained by the modification of atmospheric variables (reacting to both SW and LW perturbation). Particularly, the lower troposphere associated to a strong SW surface cooling (discussed further in section) is likely to offset the increase of LW emissivity when dust are present. Consistently, at night time, this SW cooling is no more efficient and dust layers tend to increase the downward LW radiation and warm the surface. Locally, the modification of surface incoming LW can become even more complicated if cloud cover is modified by the dust atmospheric perturbation.

Before presenting the effect of mineral dust on the top of atmosphere (Sect. 5.3) outgoing radiations, we discuss here the impact of the net (shortwave plus longwave) surface dimming on surface temperature and surface energy budget, by using sensible

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heat fluxes measurements.

5.2 Effect of dust surface dimming on the surface temperature and sensible heat fluxes

Changes in air surface temperatures due to dust surface dimming, ΔT , has been calculated as the difference between the surface temperature simulated with and without dust. Results obtained for the 9, 10, 11 and 12 March (at 12:00 UTC) are reported in the Fig. 7, indicating a significant decrease in the air surface temperature when dusts are included in simulations. Comparisons between the geographical patterns of the air surface temperature change and dust AOD distribution (Fig. 1) clearly show that the strongest decrease are related to highest dust AOD. As observed in the Fig. 7, the change in air surface temperature can be locally significant, as over Northern Benin (with a mean change of -2.8°C , for the 9 to 12 March period) or over Western Nigeria (mean of -3.7°C , for the 9 to 12 March period). At Niamey, Slingo et al. (2006) reported a drop in the daytime maximum temperature of about 10°C few days after the passage of the dust outbreak. At Djougou, we have observed a decrease of the air surface temperature up to 2°C (at noon), associated with an increase of the temperature during night time (up to 2°C at 04:00 a.m.) due to the longwave trapping by dust and re-emission. In our simulations, dust aerosol induces a surface temperature decrease of about 0.68°C (at noon) averaged over a $09^{\circ}-17^{\circ}\text{N}$, $10^{\circ}\text{W}-20^{\circ}\text{E}$ region, for 9–12 March period. This cooling could have a significant impact on the lower troposphere dynamics. This particular point will be discussed in the last part of this study (5.5).

As mentioned in the introduction, one of the interests here concerns the effect of the dust surface dimming on surface energy budget. For that purpose, we have used MesoNH simulations together with experimental observations of sensible heat fluxes (H) performed at Djougou. This approach is quite important because of the surface balance between radiation, evaporation (latent heat flux from the surface to the atmosphere) and sensible heat flux. One or all of these components will decrease to compensate for the reduction of total (direct plus diffuse) surface solar radiation. At

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this time and excepted the work of Grini et al. (2006), few work have documented the possible effect of dust particles on the surface energy budget over the Western Africa region. Most of all, no studies have coupled modelling exercises with observational dataset over the region.

As for the surface temperature, the impact of dust on sensible heat fluxes has been determined by the difference between simulated HDUST and HNONE, which are the surface sensible heat fluxes in interactive dust simulations and in the control run, respectively. Our results are reported in the Fig. 8, showing ΔH for the 9, 10, 11 and 12 March (at noon), underlying intense decreases, with maxima of 200 W/m^2 corresponding to higher dust AOD. Over Benin, simulations indicate a decrease of about $\sim 100\text{--}150 \text{ W/m}^2$, what is conjointly observed from experimental measurements at Djougou (Fig. 9), showing that the dust storm dropped by around $100\text{--}150 \text{ W/m}^2$ the surface sensible fluxes. As an example of results, H is changing from 270 to 150 W/m^2 , for the 11 and 13 March, respectively (at noon).

For comparisons, Fan et al. (2008) reported a decrease of about $20\text{--}30 \text{ W m}^{-2}$ in sensible heat fluxes in case of considerably lower AOD (0.27 at 550 nm). Jiang and Feingold (2006) indicated larger effects on surface heat fluxes (sum of sensible and latent heat fluxes), with a decrease of about 100 W/m^2 (at noon) for AOD around 1 . For biomass burning aerosols, Feingold et al. (2005) showed a reduction of 60 W/m^2 in the sensible heat fluxes (AOD of 0.6). It should be noted that these results were obtained from modelling exercises and not reinforced by experimental observations, as proposed in this work.

5.3 Dust effect on the top of atmosphere outgoing radiation

As mentioned in the Sect. 5.1, we discuss here about the effect of dust on Top Of Atmosphere (TOA) outgoing radiation, both in shortwave and longwave spectral ranges. Similarly, we characterise this effect by calculating the flux difference: $\text{TOA_SW} = (\text{SW_UP})_{\text{NONE}} - (\text{SW_UP})_{\text{DUST}}$, where SW_UP is the upward shortwave radiative fluxes (in W/m^2). Once again, with this relation, a negative sign of TOA_SW (or

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TOA_LW) implies an overall cooling effect at TOA and a positive one a heating effect. As mentioned by Roger et al. (2006), we consider that the aerosol radiative effect at TOA is defined with an accuracy of 20%.

5.3.1 Solar spectral range

5 Figure 10 displays the difference in upward shortwave fluxes, with and without aerosols, for the 9, 10, 11 and 12 March (at noon). Our simulations reveal a significant north-south gradient over West Africa. Two different zones appear, with a first one (16° – 21° N/ 10° E– 20° W) characterized by simulated positive values, with a regional mean of $+5.3$ W/m² and local maximums reaching up to $+100$ W/m². The second one (07° –
10 14° N/ 1° E– 15° W) is characterized by a mean regional of -6.6 W/m², with maximums reaching up to -100 W/m², what is consistent with values reported by Slingo et al. (2006) who indicated that reflected fluxes at the TOA rose by 100 W/m² at Niamey. Our simulation for 10 March over Senegal is also in agreement with the instantaneous TOA forcing calculated by Derimian et al. (2008) in M'Bour, which range from about
15 -70 to -90 W/m² (at noon). In addition, Milton et al. (2008) reported values comprised between -10 and -80 W/m² over the Sahel region.

This result implies two opposite effects with a heating and cooling of the “earth-atmosphere” system over different zones. Such opposite effect is mainly due to decreasing North-South surface albedo gradient. Concerning the second one (07° –
20 14° N/ 1° E– 15° W), the combined aerosol-surface system reflects more solar radiations back to space than the surface associated with a clean atmosphere. Similar results are obviously obtained over oceanic regions, expected over the Gulf of Guinea. Over the first zone (14° – 21° N/ 1° E– 15° W), the presence of dust above high reflectivity areas reduces upward fluxes compared to a surface of a specific albedo alone. Over such
25 regions, mineral aerosols warm the “earth-atmosphere” system.

Over the Gulf of Guinea in Nigeria, a large positive TOA effect is obtained for the 10 and 12 March, with values ranging from $+100$ to 200 W/m². As reported in the Sect. 5.1.1, this effect is due to changes in the cloud cover. In case of simulations

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including dust aerosols, low clouds cover is decreasing over the gulf of Guinea, leading to lesser solar fluxes reflected back to space.

In term of regional average (Table 1), the shortwave TOA radiative effect is negative with minimum values (at noon) of -1.1 W/m^2 for the 9 March, when most of the mineral dust is located in majority north of the domain (i.e., above high surface albedo). The maximum regional mean (-24 W/m^2) is simulated for the 11 March, when dust aerosols are advected south of the domain above low surface albedos, associated with largest mean dust AOD (1.23). To summarize and whatever the day studied, dust particles are always shown to cool the “earth-atmosphere” system in the solar range (Table 1).

5.3.2 Infrared spectral range

Dust effect on TOA longwave radiations (TOA.LW) are reported in the Fig. 11, for 9, 10, 11 and 12 March (at noon). Results indicate that dust aerosols produce a positive effect over the entire West Africa region, with instantaneous increases comprised between $+20$ and $+50 \text{ W/m}^2$. At Niamey, Slingo et al. (2006) derived the Outgoing Longwave Radiation (ORL) from the GERB broadband radiometer and show large signals during this dust outbreak, with a significant decrease by about 30 W/m^2 at midday, which is consistent with the analysis of previous dust storms (Haywood et al., 2005).

At the regional scale (summarized in the Table 1), our simulations indicate a mean TOA.LW comprised between $+3.3$ and $+6.7 \text{ W/m}^2$. Although the net effect (shortwave+longwave) still remains cooling at TOA for most of cases, the inclusion of longwave effect in radiative simulations offsets a part of the shortwave cooling and can even produce net positive impact at TOA, i.e. 9 March (net effect of $+2.2 \text{ W/m}^2$, Table 1). For this specific day, it should be emphasized that most of the aerosols are located north of the domain above high surface albedo. As mentioned in Sect. 5.4.1, this decreases TOA.SW cooling and increasing TOA.LW effect due to strong surface temperature.

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5.4 Effect of the dust on heating rate profiles and convective activity

5.4.1 Solar heating rates

Figure 12 reported the difference between HR_SW simulated by using two parallel simulations for the 9, 10, 11 and 12 March (at noon). First, it is seen that dust HR_SW are linked with the vertical structure of the dust plumes. Dust aerosols induce a significant perturbation in solar heating rates ranging from +4.0 to +7.0 K by day and occurring in majority within the 2–4 km dusty layer. The absorption within the dusty layer, together with the scattering, cuts down the incoming solar flux thus decreasing the heating rates of the air below the dust layer (Fig. 12).

For comparisons, Mohalfi et al. (1998) reported vertical distributions of shortwave heating rates (at noon) due to dust aerosols only over the Saudi desert as function of AOD. Results indicated maximum heating of 3.0, 4.5 and 6.5 K by day, corresponding to AOD of 0.5, 1.5 and 3, respectively. Kim et al. (2004) reported instantaneous vertical profiles of shortwave heating rates obtained onboard the C-130 research flight during Asian dust events, showing values comprised between 2 and 5 K by day in the dusty layer. Although obtained for a different site, these estimations are found to be coherent with HR_SW as simulated by MesoNH.

5.4.2 Infrared heating rates

Figure 13 shows estimations of HR_LW due to dust particles, for the 9, 10, 11 and 12 March (at noon). Results show significant LW cooling below the dust layer for each day, with values ranging from –6 to –16 K by day. This effect is quite different to those reporting found in the literature. As an example, Mohalfi et al. (1998) reported –1 K by day over Saudi Arabia and Kim et al. (2004) reported HR_LW values ranging from –2 to –4 K by day in the dusty layer over Asian region.

In our case, the large LW cooling below the dusty layer is due to the complete online calculation of its term in MesoNH. Here and as mentioned in Sect. 5.1.1, the presence

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of dust strongly decreases the solar energy accumulated by the surface, cooling it (Table 1), and decreasing by this way upward surface LW fluxes potentially absorbed by the layer below dust (0–2 km), which in turn is logically cooled. This result underlines the important role of the surface shortwave dimming on the infrared budget. Over the 16°–21° N/10° E–20° W region, the mean surface emission is reduced by 21.4 W/m². This effect is well observed at Niamey by Slingo et al. (2006) who showed that the surface emission drops locally by over 100 W/m² during this dust event, corresponding to instantaneous surface cooling of 13°C.

This result of simulations emphasizes with simulated HR_LW for the 12 March over the Gulf of Guinea (where large AOD~2.5 are simulated, see Tulet et al. (2008)) since the ocean temperature does not respond on the timescale of a forecast. Here, the LW effect is different and displays a warming below the base of the dust layer, due to the increase of downwelling LW flux, which enhances the atmospheric absorption in the lower atmosphere. For comparisons, our results are found to be consistent with those reported by Zhu et al. (2007), who reported similar dust HR_LW over three different oceanic regions.

5.4.3 Impact of dust radiative effect on the convective activity

In order to investigate the possible effect of dust on the convective activity, we have used the mean regional convective available potential energy (CAPE) as simulated by MesoNH for 9, 10, 11 and 12 March (at noon). Our results are reported in the Fig. 14. It should be recalled that in a case of convectively unstable atmosphere, CAPE is greater than zero and increases when the temperature or moisture in the boundary layer increases (or the upper layer temperature decreases). In a stable atmosphere, CAPE is zero and changes of temperature or moisture do not affect CAPE unless the changes are large enough to make the atmosphere unstable.

In Fig. 14, we have reported CAPE vertical profiles of the difference between MesoNH simulations including or not dust aerosols. Results of simulations show a remarkable decrease between the surface and 2 km height with a value reaching up to

–350 J by kg at the surface. Although only obtained through modelling exercises and not validated by in-situ observations at this time, this indicates the potential decrease of the convection processes due to dust radiative effects over the West African region.

This work seems to confirm results previously obtained by Grini et al. (2006) over the Sahel region by using MesoNH and simulations conducted by Jiang and Feingold (2006), Fan et al. (2008) or Wendisch et al. (2008), who show that the reduction of surface incoming radiation due to aerosols produce a weaker convective atmosphere.

6 Conclusions

A study of the dust aerosol impacts on the radiative effect, surface energy budget (surface temperature and sensible heat fluxes), heating rate profiles and convective activity over West Africa is presented. The study is based on numerical MesoNH simulations coupled with experimental AMMA observations obtained at Djougou (Benin) during an intense dust event (9–13 March). Two numerical simulations have been conducted: one including the radiative effects of dust aerosols and another without. The major results are the following:

- The presence of dust particles induces a large instantaneous (at noon) reduction of surface incoming shortwave radiation (with a regional mean of -160 W/m^2) over a large part of West Africa ($09^\circ-17^\circ \text{ N}$, $10^\circ \text{ W}-20^\circ \text{ E}$) during the 9 to 12 March period.
- The surface dimming resulting from the presence of dust is shown to cause important reductions in both the surface temperature (up to 4° C at midday over regions where high AOD occur) and in sensible heat fluxes (up to 100 W/m^2). These modelling results are shown to be consistent with the experimental observations performed at Northern Benin.
- The net effect (shortwave+longwave) at TOA is dominated by the SW one although the mean positive LW effect (regional mean of $+5 \text{ W/m}^2$) compensates

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about half of the mean negative SW one (regional mean of -13.5W/m^2). As a result, the presence of dust plumes cools the earth-atmosphere system.

- The atmospheric heating rate changes induced by dust show large SW heating (at noon) within the dusty layer (2–4 km) due to strong absorption of the incoming solar radiation, with the maximum heating comprised between +4.0 and +7.0 K by day. In the LW, a significant cooling is obtained below the dust layer (0–1 km), with values ranging from –6 to –16 K by day. This effect is due to the above discussed surface temperature decrease which strongly reduces surface LW emissions and hence warming of lower troposphere layers.
- The surface cooling and reduction of energy fluxes associated with aerosol heating in the dust layer results, in our a case, in a reduction of convective activity, with a mean decrease of CAPE reaching about 200J kg^{-1} between the surface and 2 km height.
- Further studies should be conducted for investigating the impact of mineral dust on the stratocumulus clouds cover over the Gulf of Guinea.

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Table 1. Instantaneous (noon) regional mean (09° – 17° N, 10° W– 20° E) of TOA and SRF dust radiative effect (in W/m^2) simulated in shortwave (SW) and longwave (LW) regions.

	Total		Total		Total		Total		Total	
	SW	LW	SW	LW	SW	LW	SW	LW	SW	LW
TOA	-1.1	+3.3	-8.1	+5.7	-23.9	+6.7	-10.6	+4.0	-11.4	+3.7
	+2.2		-2.4		-17.2		-6.6		-7.7	
SRF	-105.1	-3.7	-151.9	-3.0	-196.6	-2.3	-202.8	-4.0	-176.8	-5.3
	-108.8		-153.9		-198.9		-206.8		-182.1	
	9 March		10 March		11 March		12 March		13 March	
	(AOD=0.60)		(AOD=0.83)		(AOD=1.23)		(AOD=1.13)		(AOD=0.95)	

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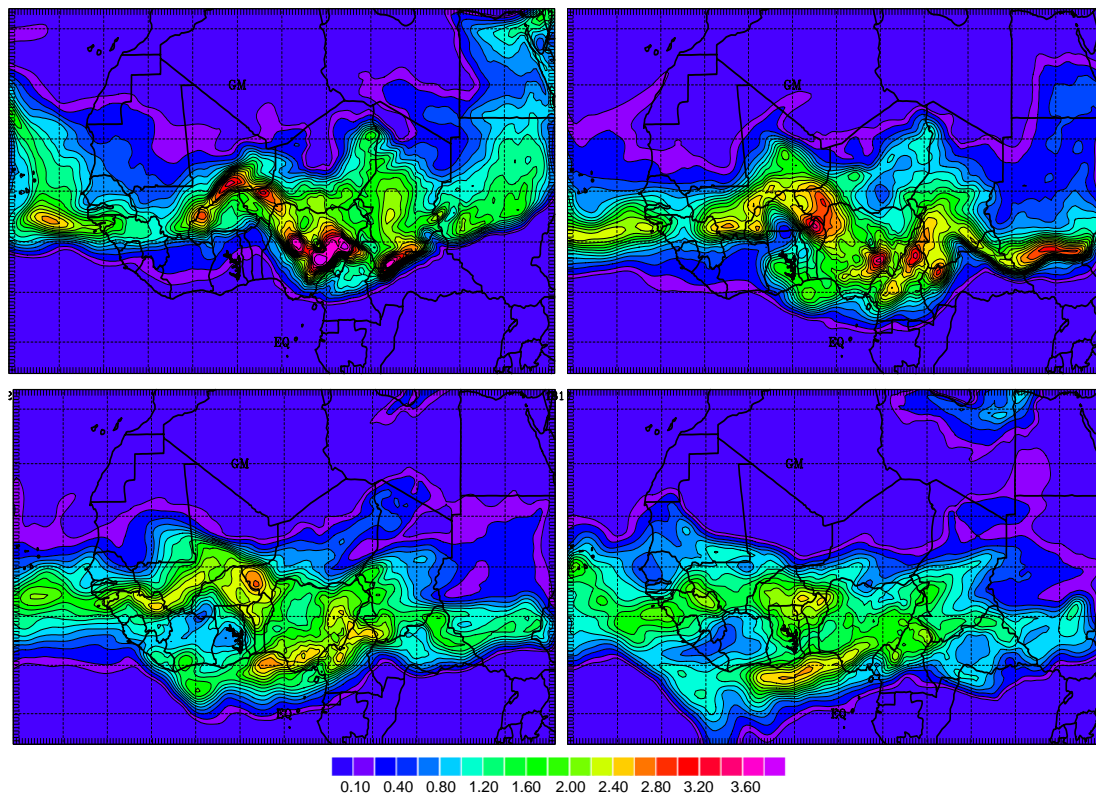


Fig. 1. Dust AOD (at 560 nm) simulated at noon by MesoNH for the 9 (left up), 10 (right up), 11 (left down) and 12 March (right down) 2006.

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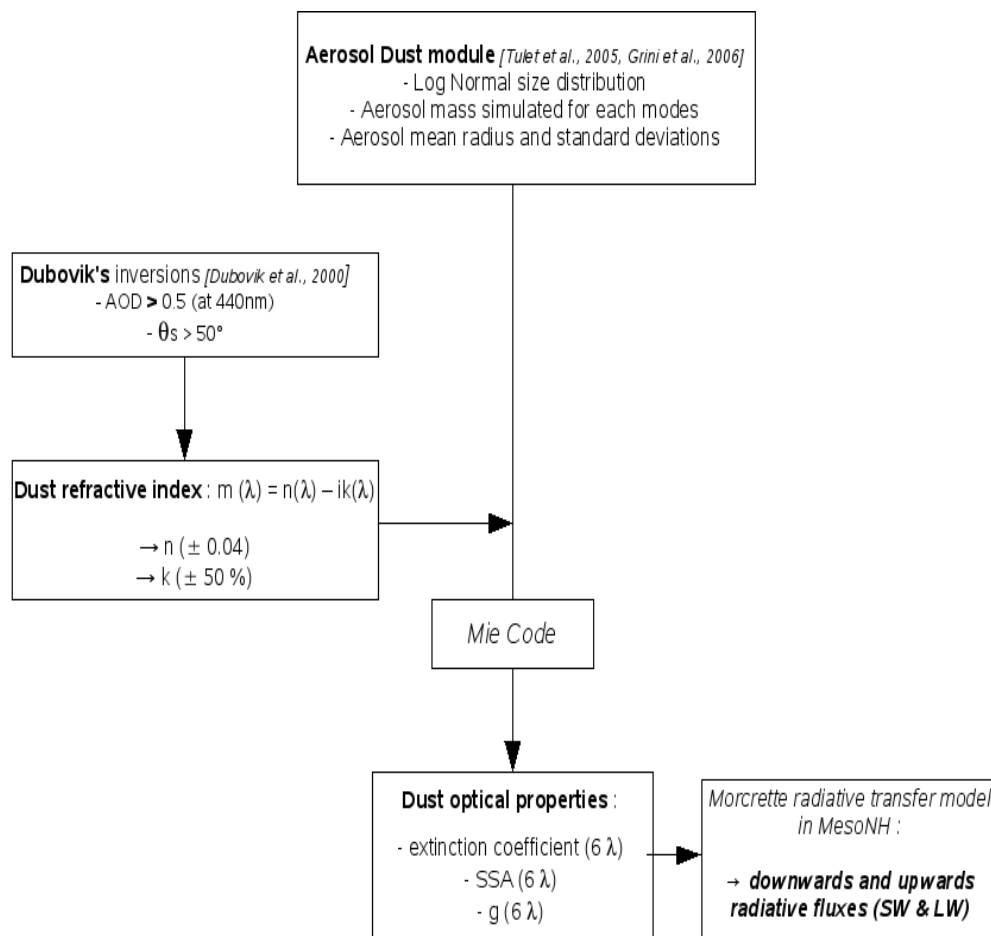


Fig. 2. Methodology used for computing the dust direct radiative effect in MesoNH.

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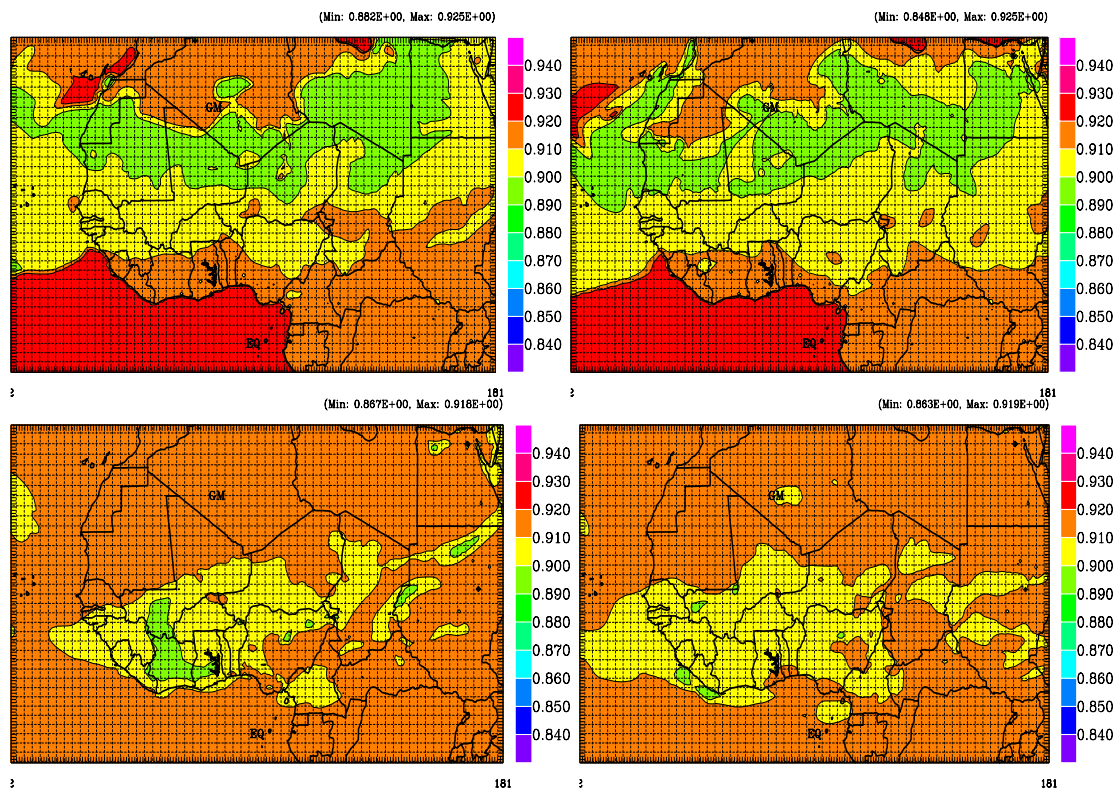


Fig. 3. Dust SSA (at 560 nm) estimated at noon by MesoNH at the surface (bottom) and 2 km height (down), for the 9 (left) and 10 (right) March.

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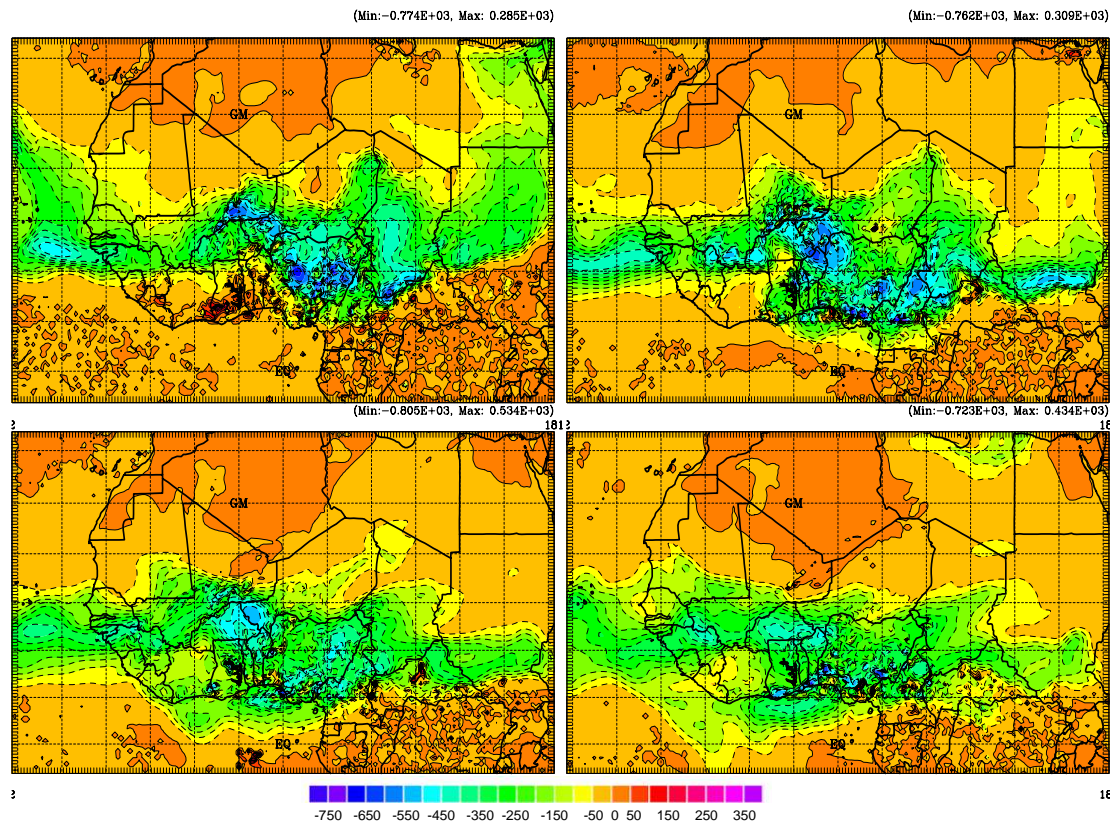


Fig. 4. Instantaneous (noon) dust effect on surface shortwave downward radiations (in W/m^2) simulated by MesoNH for the 9 (left up), 10 (right up), 11 (left down) and 12 (right down) March 2006.

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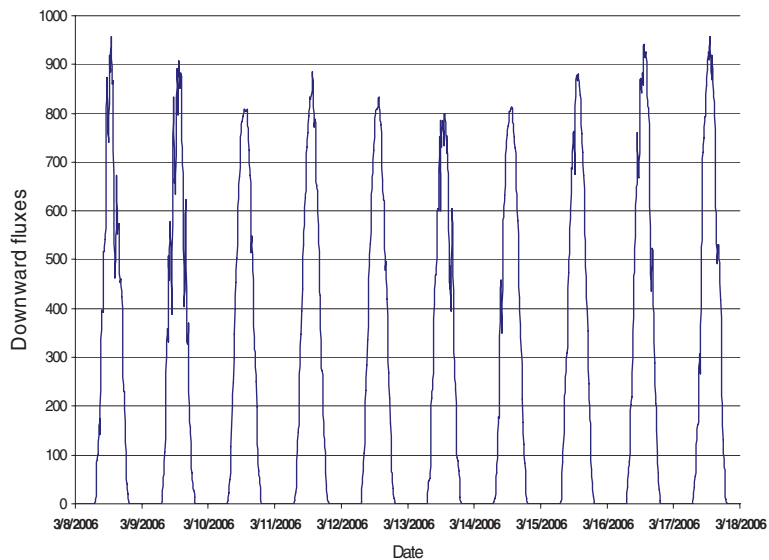


Fig. 5. Downward shortwave fluxes (in W/m^2) observed at Djougou from pyranometers (Kipp and Konen) measurements for the 8 to 18 March period.

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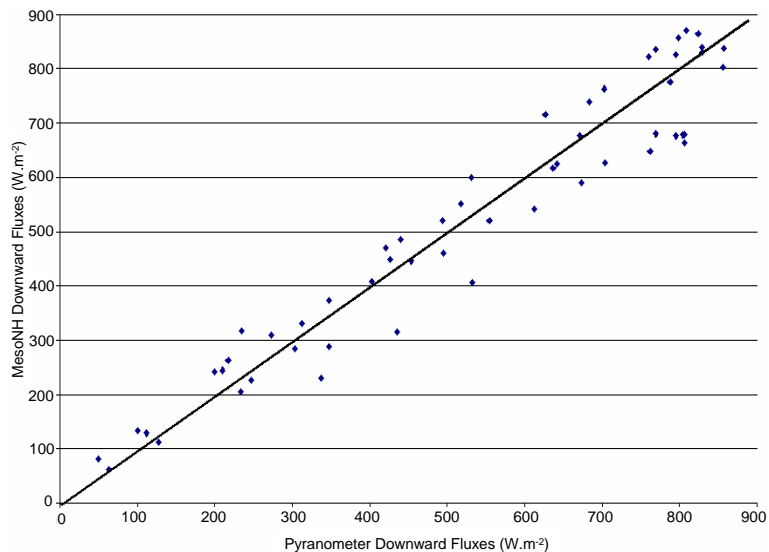


Fig. 6. Comparisons between downward shortwave fluxes simulated by MesoNH and measured at Djougou for the 10–12 period of March (pure clear-sky days). The solid line corresponds to 1:1 line.

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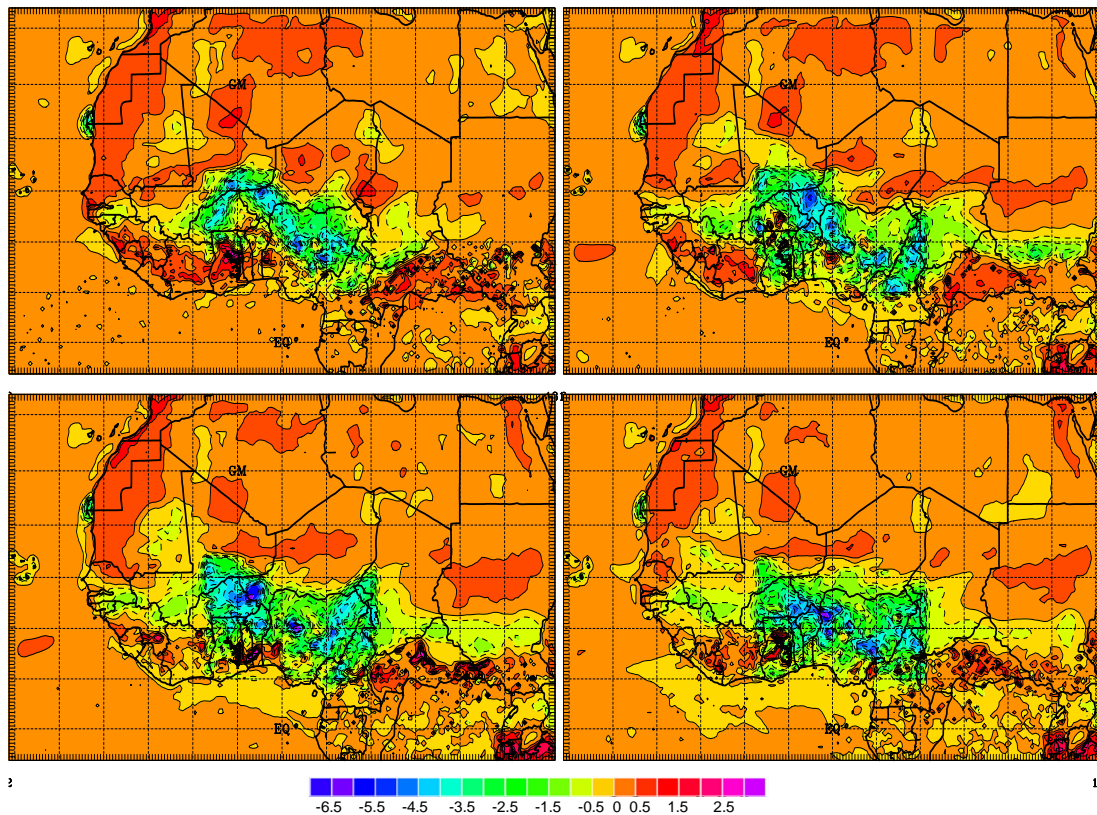


Fig. 7. Instantaneous change (at noon) in surface temperature simulated by MesoNH for the 9 (left up), 10 (right up), 11 (left down) and 12 (right down) March 2006.

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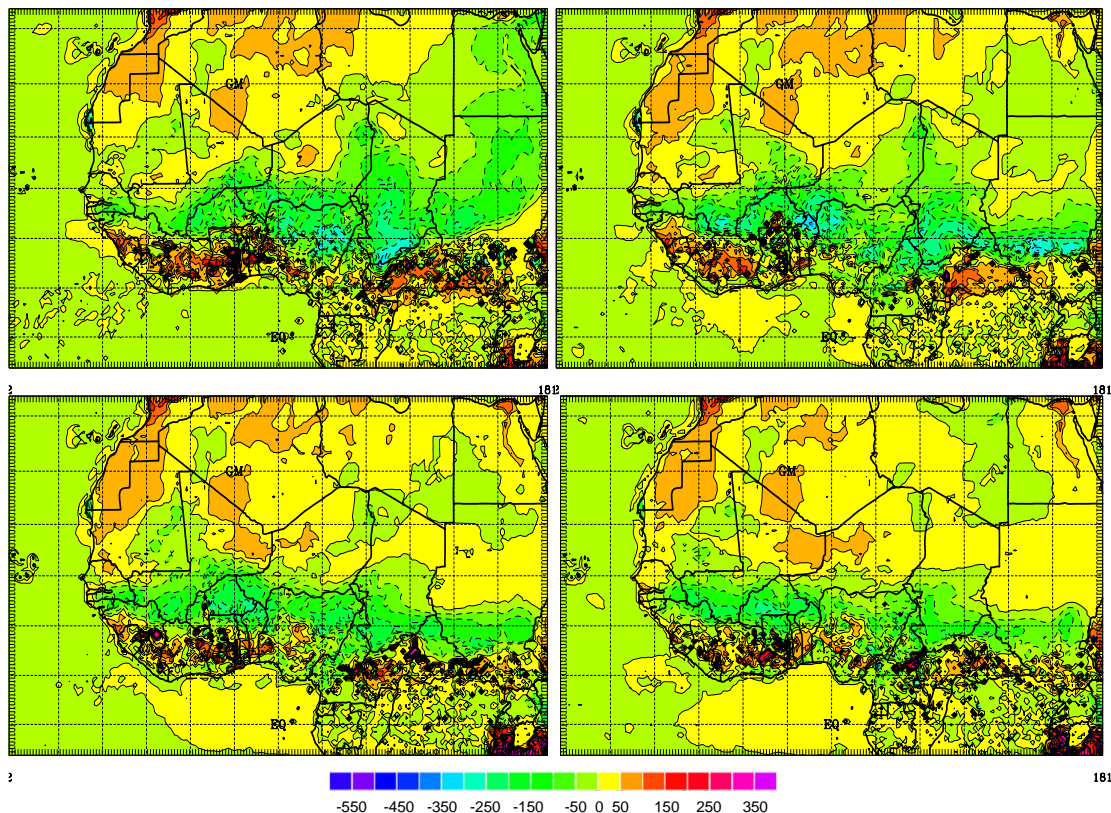


Fig. 8. Instantaneous change in surface sensible fluxes (at noon) due to dust aerosols for the 9 (left up), 10 (right up), 11 (left down) and 12 (right down) March 2006.

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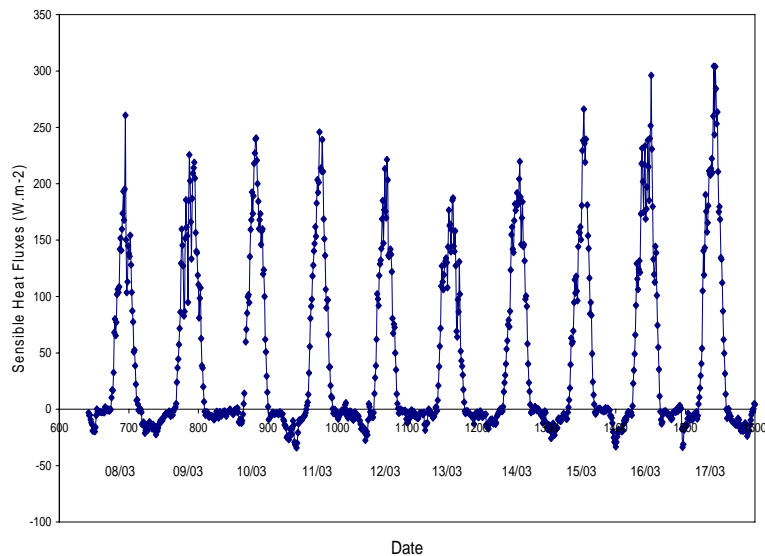


Fig. 9. Sensible heat fluxes (in W/m^2) observed at Djougou during the 8 to 17 March period.

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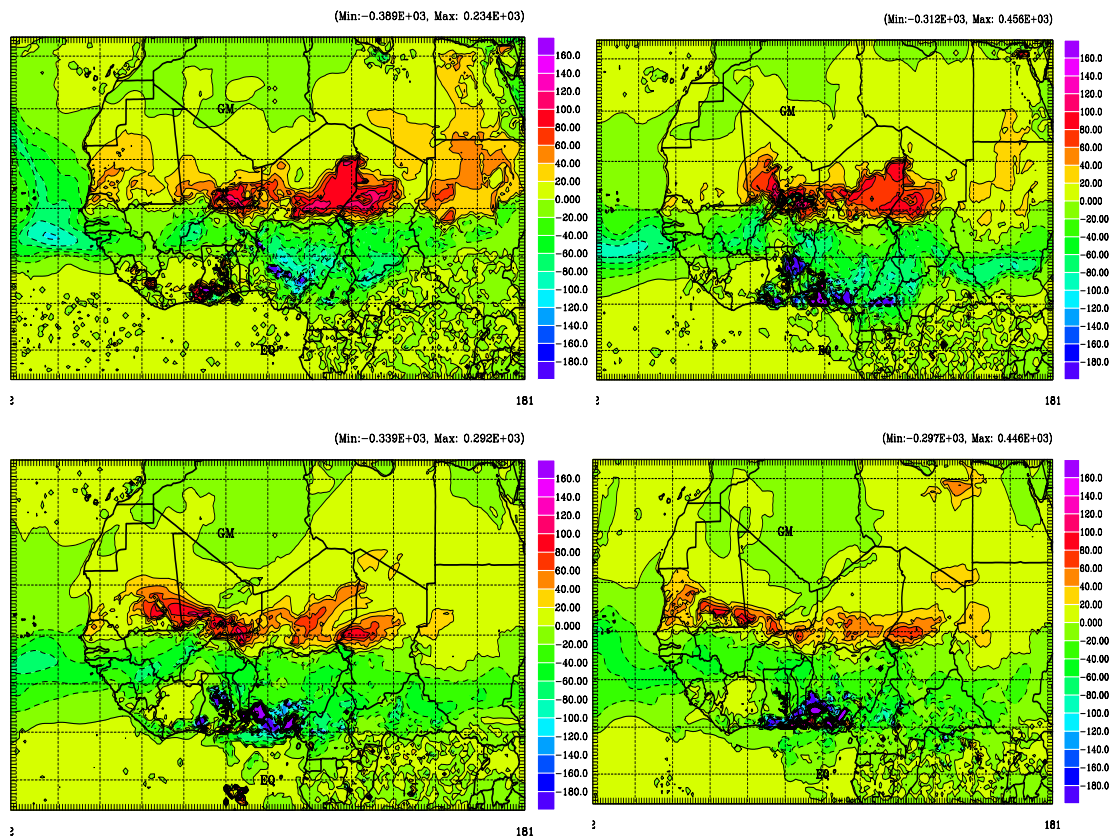


Fig. 10. Instantaneous (noon) dust TOA shortwave effect (in W/m^2) simulated by MesoNH for the 9 (left up), 10 (right up), 11 (left down) and 12 (right down) March 2006.

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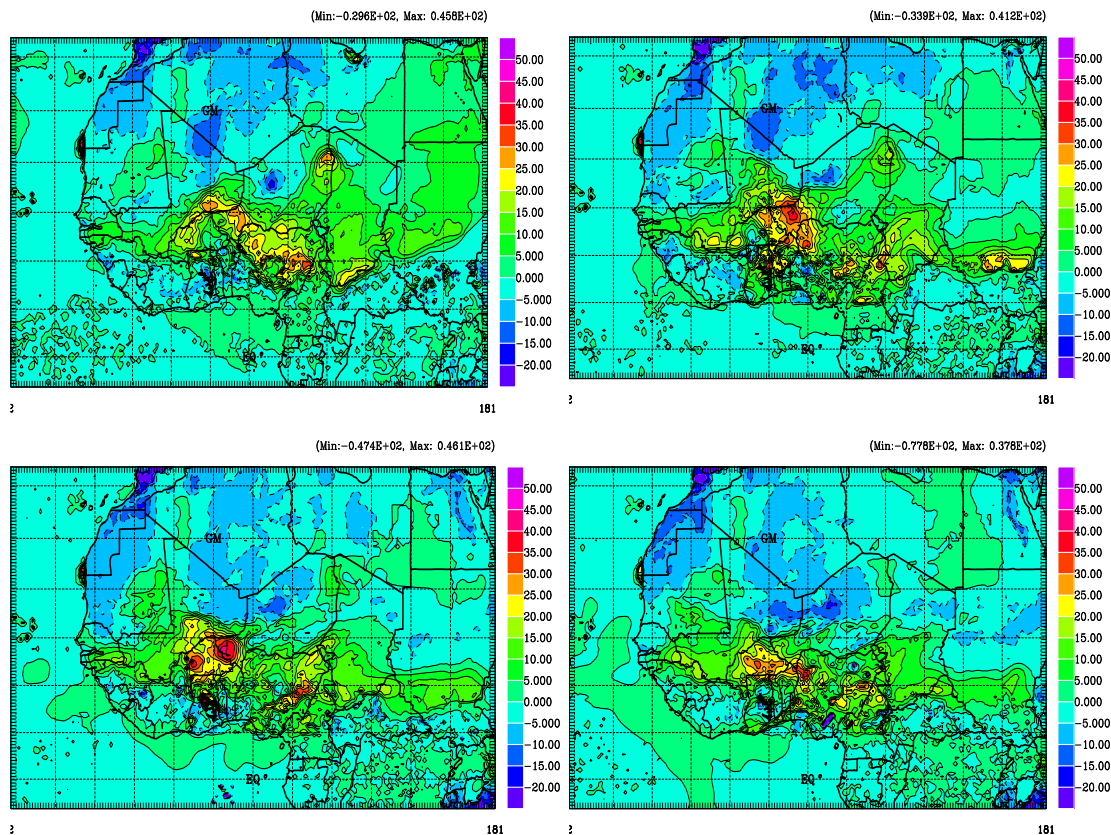


Fig. 11. Instantaneous (noon) dust TOA longwave effect (in W/m^2) simulated by MesoNH for the 9 (left up), 10 (right up), 11 (left down) and 12 (right down) March 2006.

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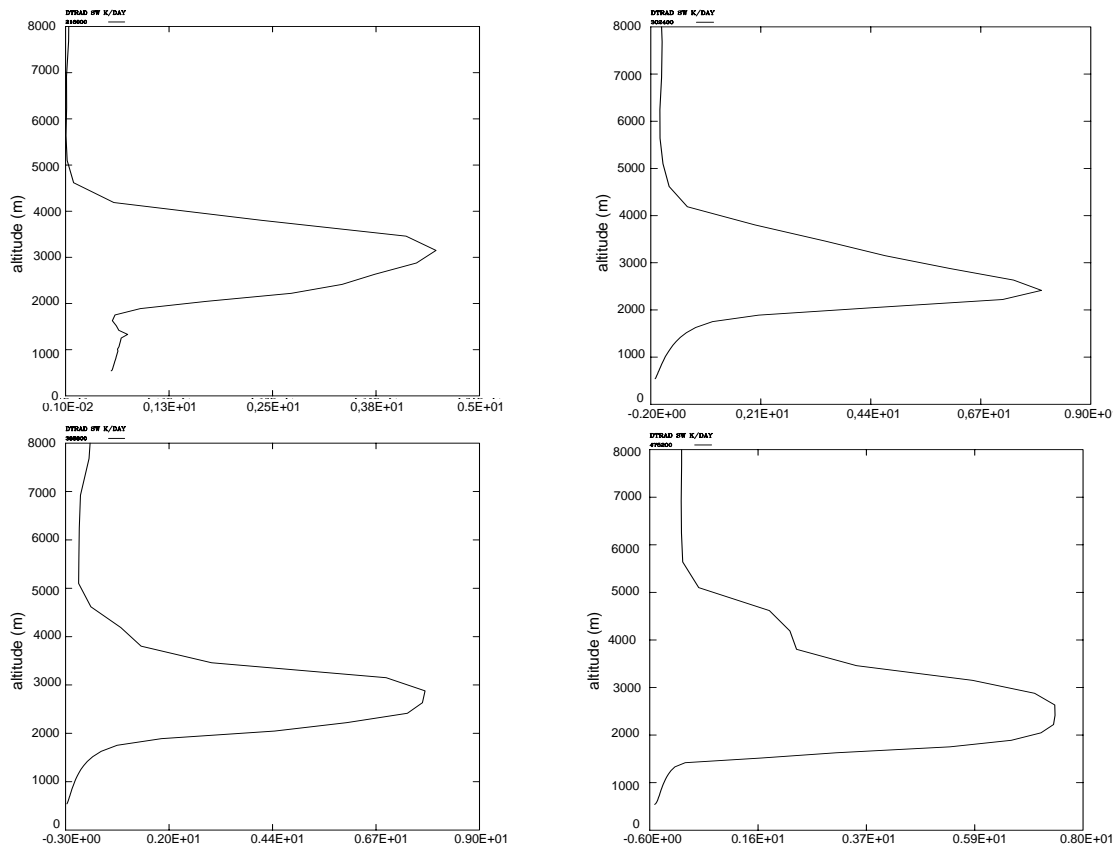


Fig. 12. Regional averaged (09° – 17° N, 10° W– 20° E) of dust aerosol-induced perturbations in the solar heating rates (in K by day), for the 9 (left up), 10 (right up), 11 (left down) and 12 (right down) March 2006.

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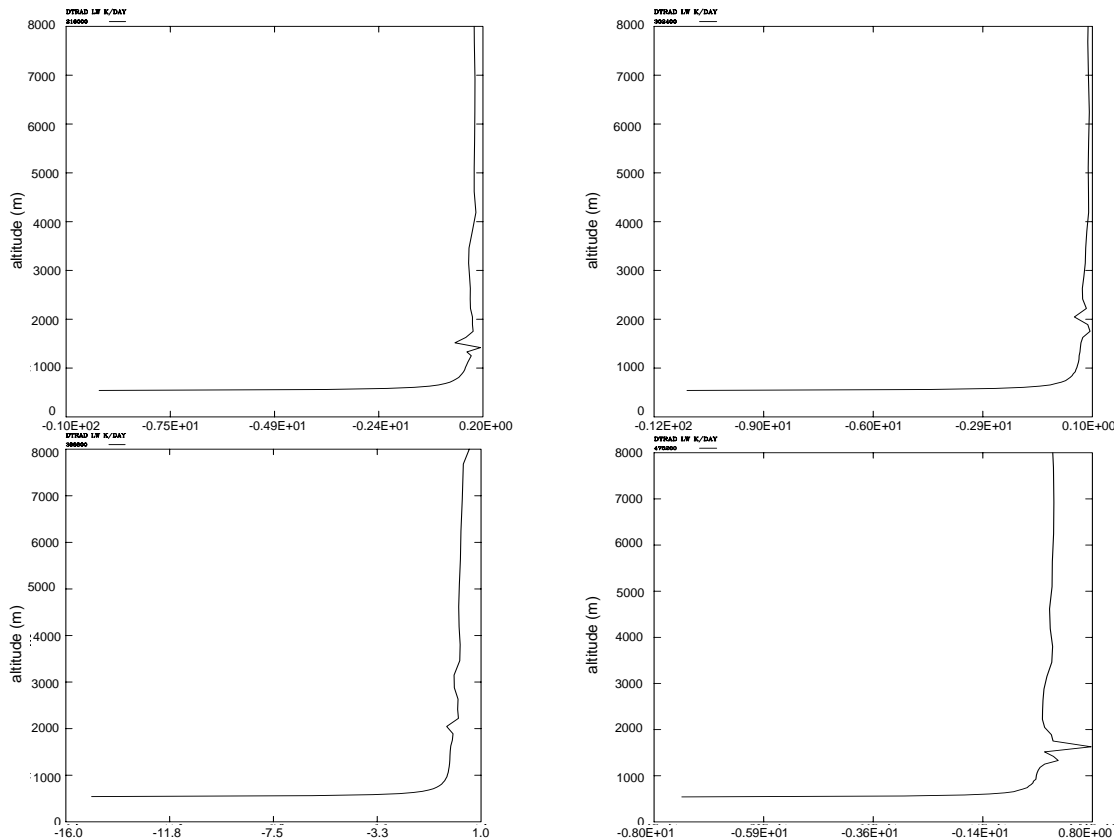


Fig. 13. Regional averaged (09° – 17° N, 10° W– 20° E) of dust aerosol-induced perturbations in the infrared heating rates (in K by day), for the 9 (left up), 10 (right up), 11 (left down) and 12 (right down) March 2006.

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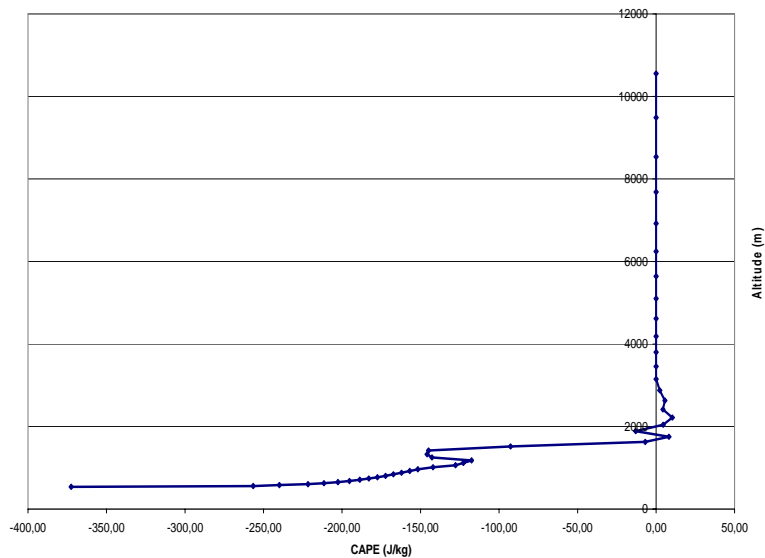


Fig. 14. Regional averaged (09° – 17° N, 10° W– 20° E) of dust aerosol-induced perturbations in CAPE (in J by kg), averaged of the 9, 10, 11 and 12 March (at noon).

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